

The influence of cloud top variability from radar measurements on 3-D radiative transfer

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Abstract

In radiative transfer simulations the simplification of cloud top structure by homogenous assumptions can breed to mistakes in comparison to realistic heterogeneous cloud top structures. This paper examines the influence of cloud top heterogeneity on the radiation at the top of the atmosphere. The use of cloud top measurements with a high temporal resolution allows to analyse small spatial cloud top heterogeneities by using the frozen turbulence assumption for the time – space conversion. Radiative observations are often based on satellite measurements, whereas small spatial structures are not considered in such treatments. A spectral analysis of the cloud top measurements showed slopes of power spectra between -1.8 and -2.0 , this values are larger then the spectra of $-5/3$ which is often applied to generate cloud field variability. The comparison of 3-D radiative transfer results from cloud fields with homogenous and heterogeneous tops have been done for a single wavelength of $0.6 \mu\text{m}$. The radiative transfer calculations result in lower albedos for heterogeneous cloud tops. The differences of albedos between heterogeneous and homogeneous cloud top decrease with increasing solar zenith angle. The influence of cloud top variability on radiances is shown. Explicitly in forward direction the reflectances for heterogeneous tops are larger, in backward direction lower. The largest difference of the mean reflectances (mean over cloud field) between homogeneous and heterogeneous cloud top is approximately 0.3, which is 30% of illumination.

1 Introduction

The importance of clouds in the climate system is indisputable, because they strongly influence the solar insolation, the most significant energy source for the climate system. Clouds are spatially highly inhomogeneous, which is determined by variations in cloud microphysics and cloud geometry. Up to now satellite measurements are not able to gauge cloud describing parameters in a spatial adequate resolution, neither

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for microphysical parameters nor for geometrical ones. But these variabilities, in the so called “subpixel” scale, strongly influence the radiative transfer. Even [Randall et al. \(2003\)](#) showed the correlation of smaller and larger scale behaviour of the atmospheric system. Especially in the field of radiation calculations in global atmospheric models [Randall et al. \(2003\)](#) adduced, that the parameterisation of the input parameter, like phase, shape and size of cloud particles but also cloud geometry is the main reason for inaccuracies of radiative transfer results.

Many studies have used stochastic cloud fields to investigate the influences of variabilities of macro- and microphysical parameters on radiative transfer. In most cases the variabilities have been attributed to variations in volume extinction coefficient whereas cloud geometry have been kept constant (e.g. [Barker and Davies, 1992](#); [Marshak et al., 1995a,b](#)). Already [Loeb et al. \(1998\)](#) and [Loeb and Coakley \(1998\)](#) revealed that the cloud top structure may also have substantial effects on the radiative transfer. So the influence of cloud variability cannot be explained by assuming variations only in cloud microphysics, keeping cloud geometry, especially cloud top height, constant. By a look at the sky everybody has already noticed that the assumptions of flat cloud bottoms or tops are truly inappropriate, even for stratiform clouds.

Therefore, the main purpose of this paper is to investigate the influences of cloud top variability on radiative transfer. This task should be done by describing the clouds as realistic as possible. But at the same time the cloud top variability has to be the only cause of differences in radiative transfer results. In this study a full 3-D radiative transfer calculation is performed by using a Monte Carlo algorithm.

To describe the variability of the cloud top no constant value, like $-5/3$ for the slope of the power spectrum, is implemented. Rather high-resolution radar and ceilometer measurements are used to derive the variability, especially from cloud top. Atmospheric parameters used in this study like wind, temperature and pressure have been recorded simultaneously at the Meteorological Observatory Lindenberg.

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2 Methodology

2.1 Simulation of cloud fields

In these study measurements of three clouds are chosen to simulate the cloud fields for radiative transfer calculations. All clouds are assumed to consist completely of liquid water. The first ice particles in super cooled clouds appear at temperatures between 263°K and 258°K ([Lamb, 2002](#)). In Table 1 the simulated cloud fields are characterized (CT = cloud top, CB = cloud base).

Clouds have been chosen to cover a great part of the natural diversity of geometrical cloud characteristic. Whereas cloud 1 is the type geometrical thick cloud with variable top, cloud 2 geometrical thick with less variable cloud top and cloud 3 represents a geometrical thin cloud with variable top. For these three measured clouds the following way of simulating cloud fields for radiative transfer calculations is performed. The transformation from time to space of cloud top heights, measured by vertical pointing radar is based on the so called “frozen turbulence assumption”, which assumes no changes of the cloud field during the measurements.

To generate the 2-D cloud top field from 1-D measurement data, the iterative amplitude adjusted Fourier transform (IAAFT) algorithm developed by [Schreiber and Schmitz \(1996, 2000\)](#) have been applied. This method is based on the application of Fourier spectra to characterize two point statistics of spatial or temporal data. Fourier methods have widely been used in previous studies for cloud modelling (e.g. [Barker and Davies, 1992](#)). With the help of the IAAFT algorithm the step from a one dimensional time series to a two dimensional data field has been done. The improvement of the IAAFT is that the simulated field and the measured time series of cloud top height are equal in power spectrum and the amplitude distribution, respectively. From a measured time series, s_n (with n is time or space) with N values, the power spectrum S_k (with k wave

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numbers) is calculated as

$$S_k^2 = \left| \sum_n s_n e^{\frac{i2\pi kn}{N}} \right|^2. \quad (1)$$

The relevant value that describes the variability of the time series, is the slope of a power law regression of the power spectrum and the corresponding wave numbers. A straight line continuation of the slope in the scope of higher frequencies is dependent on the absence of scale breaks in the power spectrum. Furthermore a sorted list of the measured values s_n is necessary for the IAAFT algorithm. The iteration starts with a random shuffle of s_n . The first step of the algorithm is to adjust the Fourier coefficients, the second is the adjustment of the amplitudes. To achieve the desired power spectrum the Fourier transform of the time series is calculated in each iteration. The absolute values of the coefficients are replaced by those from the measured time series, but the phases are retained. A backward transform of these coefficients would produce an amplitude distribution which is not the same as the measured one. Therefore the second step is the adjustment of the amplitude distribution, where the amplitudes are sorted and replaced by the sorted values of the original values. These two steps of the iteration have to be repeated until the power spectrum and the amplitude distribution of generated and measured values are matching in sufficient conditions.

The derivation of a 2-D variability grid from a 1-D spectrum with the assumption of isotropic statistics yields to an underestimation of the variance of the 2-D field. This means that the slope of a single row of the 2-D field is much lower than the slope of the 1-D time series. This problem is discussed in [Austin et al. \(1994\)](#), and they propose to use

$$\gamma = \beta + 1 \quad (2)$$

where β is the slope of the 1-D spectrum of the measured time series and γ is the spectrum that produces a 2-D field consisting of the 1-D value β .

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So, cloud top fields were generated consisting of the same power spectrum and amplitude distribution as the 1-D measured time series of cloud top height.

The vertical resolution due to the measurements is assigned to the 2-D field by the IAAFT. To get a higher vertical resolution of cloud tops for the 3-D cloud field, a linear interpolation of the cumulative amplitude distribution is performed. So the second step of the IAAFT, the amplitude adjustment, is done using a refined amplitude distribution.

The vertical dimension of the cloud field is characterised by a subadiabatic liquid water content (LWC) and an adiabatic profile of the effective radius. The LWC profile is based on the study of [Chin et al. \(2000\)](#). In this study a weighting function is applied to describe the subadiabatic character of the profile. This weighting function is given by

$$f(\hat{z}) = \exp(-\alpha \cdot \hat{z}^\beta) \quad (3)$$

where \hat{z} is the scaled height within the cloud and α and β are positive constants. In the study, [Chin et al. \(2000\)](#) differ between two types of weighting functions: one is related to subadiabatic conditions involving cloud top entrainment alone and the other considers both cloud top entrainment and drizzle effects. To assure the validity of Mie theory for calculation of optical properties the first type was chosen, with the parameterisation of $\alpha = 1.375$ and $\beta = 4$. The value of α is recommended by [Chin et al. \(2000\)](#) and with $\beta = 4$ a strong cloud top entrainment is simulated. The adiabatic LWC-profile and the weighting function are calculated from cloud base to the highest cloud top. Then the accordant values for the discretised heights are interpolated and allocated to the overall cloud level.

The adiabatic profile of the effective radius is calculated using the study of [Brenquier et al. \(2000\)](#). The way of calculation is the following,

$$LWC_{ad}(h) = C_w \cdot h, \quad (4)$$

$$r_{v,ad}(h) = (A \cdot h)^{\frac{1}{3}} \cdot N_{ad}^{-\frac{1}{3}}, \quad (5)$$

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with: $A = \frac{C_w}{\frac{4}{3}\pi\rho_w}$,

$$r_{e_{ad}}(h) = k^{-\frac{1}{3}} \cdot r_{v_{ad}} \quad (6)$$

$$= (A \cdot h)^{\frac{1}{3}} \cdot (k \cdot N_{ad})^{-\frac{1}{3}} \quad \text{and}$$

$$r_{s_{ad}} = k^{\frac{1}{6}} \cdot r_{v_{ad}} \quad (7)$$

5 Wherein C_w is the moist adiabatic condensate coefficient, h is the altitude above cloud base, ρ_w the liquid water density, r_v the mean volume radius, r_e the droplet effective radius and r_s the mean surface radius of the droplet size distribution. The parameter k relates r_v and r_e and N is the droplet number concentration in the cloud. The subscript “ad” for N , r_e , r_v and r_s refers to the adiabatic values. According to Brenguier et al. (2000) k is set to 0.67 for continental air masses and N_{ad} is 250 cm^{-3} representing polluted air.

This combination of a non-adiabatic LWC and an adiabatic profile of the effective radius is corresponding to the term of “inhomogeneous mixing”, mentioned in Baker et al. (1980). This mixing scheme takes place if the time of evaporation of a droplet with radius r is smaller then the time for the complete mixing process in the layer. In this case all droplet-radii in the volume affected by entrainment completely evaporate.

Figure 1 shows measured time series of cloud top height, a slice of geometrical properties of the simulated cloud field and profiles of LWC and effective radius. Figure 2 illustrates the three dimensional cloud field based on these data. As the counterpart to the cloud field with heterogeneous top a field with homogeneous cloud top has been generated using the mean cloud top height of the measured data.

2.2 Monte Carlo simulations

Monte Carlo simulations are performed with the model MC-UNIK, described in Macke et al. (1999). The model assumes periodic boundary conditions in x and y-direction.

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Each simulation runs with 10^6 photons, which are uniformly released at the top of the domain.

The solar zenith angle is set to 0° , 30° and 60° , the solar azimuth angle is constant at 0° ; observation angles are 0° , 30° and 60° for zenith angle and 0° , 60° , 120° and 180° for azimuth angle, respectively. Cloud optical properties, like volume extinction coefficient, single scattering albedo and phase function are calculated by Mie theorie for a wavelength of $0.6 \mu\text{m}$ assuming a modified gamma distribution for cloud droplet sizes.

Outside the cloudy regions rayleigh scattering has been applied, inside the cloud rayleigh and mie scattering are considered. The absorption of molecules has been neglected.

The surface albedo is examined as lambertian reflection. The value is calculated from a bidirectional reflectance distribution function (BRDF) for pasture land. This albedo is also known as “white-sky” albedo (Lucht, 2000). The parameterisation of the BRDF for pasture land is taken from Rahman et al. (1993).

3 Results

An advantage of this study is the use of cloud top variability from radar data. In many studies, power spectra are represented via their slopes in log – log plots calculated by least squares linear regression (assuming power law behaviour). Even Loeb et al. (1998) assumed the widely used slope of $-5/3$ to generate cloud top fields. The analysis of the measured time series revealed that the slopes with values of 1.8–2.0 are always larger then $5/3$ (Fig. 3). The conclusion is that lower frequencies and with this the spatial (or temporal) higher variabilities have more portion on the description of the cloud top variance. So in this study the spectra of the measured cloud top data were used to generate cloud top fields.

The focus of this study is the comparison of the radiative transfer results regarding the differences between clouds with homogeneous tops and those with heteroge-

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neous ones. The albedo as the mean value over the whole cloud field provides a first overview. Reflectances in several directions deliver more insight. Figure 4 shows the calculated albedo values for cloud 1, whereas Table 2 summarises the albedo results for all three clouds. The calculated difference is defined as heterogeneous albedo minus homogeneous one. It can be seen that the albedo for heterogeneous cloud top is lower in most cases and increases with increasing solar zenith angle (θ_{sun}). The largest difference is about 1.1%.

Figure 4 indicates that besides cloud top variability also illumination angle (here only changes in solar zenith angles) influences the albedo.

There are higher albedo values with increasing solar zenith angle (θ_{sun}), whereas the differences between homogeneous and heterogeneous cloud top are decreasing. One reason of the albedo increase with increasing θ_{sun} is according to the one-dimensional radiative transfer effect (Varnai and Davies, 1999). The reason for this increase is that cloud particles scatter light preferable in forward direction, whereby for overhead sun the solar radiation penetrates deeper into the cloud. This behaviour is well illustrated by the comparison of the photons penetration depth of the different illumination angles (Figs. 5, 6 for cloud 1).

Furthermore cloud fields tend to appear more homogeneous from oblique direction than from above, which enhances this increase (Varnai and Davies, 1999). The cause of the deeper penetration at heterogeneous cloud tops is the larger surface which leads to more transitions between cloudy parts and non-cloudy ones (Varnai and Davies, 1999). This added transport into the cloud is also pictured by transmission and absorption (Table 3). The simulated albedo values for cloud 2 and 3 are similar, increasing albedo values and decreasing differences between homogeneous and heterogeneous tops with increasing θ_{sun} . The significant difference between homogeneous and heterogeneous cloud top at $\theta_{\text{sun}} = 0^\circ$ is 0.2% for cloud 1, cloud 2 with less variability shows only a difference of 0.1% and the thin and variable cloud 3 shows the largest difference of 1.1%. The high transmission of cloud 3 with simultaneous low absorption is caused by the short vertical expansion of this cloud.

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The results mentioned above denote already some aspects of the influence that cloud top variability has on radiative transfer, which is first the lower albedo of heterogeneous cloud top and second the larger penetration depth. Now the effects on reflectances are focused. Reflectances are calculated for nine observation angles, for 30° and 60° zenith with changes in azimuth of 0° , 60° , 120° and 180° respectively and the direction of 0° zenith and 0° azimuth.

The reflectances of this observation angles are simulated for the three solar zenith angles of 0° (Fig. 7), 30° and 60° zenith and 0° azimuth.

Figure 7 shows the calculated reflectances for cloud 1 as mean values (mean over cloud field) with corresponding minimum and maximum values. The azimuth angle of illumination is 0° , so the azimuth observation angle of 0° is the backward direction relative to illumination, 180° is forward and 60° and 120° are sideways, respectively. The reflectance is defined as the ratio of reflected to incident radiation. The variability of the reflectances for homogeneous cloud tops shown in Figure 7, are the result of the uncertainty of the Monte Carlo model. These uncertainty is determined by the random nature of the Monte Carlo model and though using the local estimation approach with an obviously to low number of simulated photons.

Figures 8 and 9 show the calculated differences, defined as heterogeneous reflectance minus homogeneous one. The maximum difference of the mean reflectances between homogeneous and heterogeneous cloud top is approximately 0.3, which is 30% of illumination. The largest differences appear in forward and backward direction relative to illumination direction, whereas the differences have a negative maximum in backward direction (homogeneous > heterogeneous) and a positive one in forward direction.

It can be seen that the largest negative difference is found in backward direction when the zenith angle of illumination and observation are equal. The maximum positive difference is found in forward direction for equal zenith angles of illumination and observation.

The preferred forward direction can be explained by the forward peak of the Mie

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phase function. The probability that a photon turns around in the backward direction, is very low compared to proceeding in forward direction. Mentioned above, the cloud top heterogeneity causes a higher transport into cloudy regions, which leads to more intense scattering. The probability of a complete photon turn is much lower for clouds with heterogeneous tops. Therefore the reflectance in backward direction is lower for clouds with heterogeneous tops than for homogeneous ones, which leads to the negative differences shown in the Fig. 8. The largest reflectances appear in forward direction (azimuth angle of 180°) when the zenith angle of illumination and observation are equal. This behaviour is determined by two effects. One is the dominance of the forward peak of Mie scattering, that causes on one hand the escape of the photon in forward direction and on the other hand the deeper penetration for lower solar zenith angles. The deeper penetration effects that the distance the photon has to propagate back to the detector is larger for lower zenith angles than for higher ones. This longer way is the second effect and causes a higher attenuation of the photons energy.

4 Conclusions

Although earlier studies examined the influence of cloud top variability on radiation, only a few have used measured data with high resolution. Several of the applied techniques have been used in earlier studies, but not necessarily in this way. This study combines them and therefore tries to describe clouds as realistic as possible, always keeping in mind that cloud top variability has to be the only cause of differences in radiative transfer results. The present study should extend the earlier studies that deal with cloud top heterogeneity effects on radiative transfer in cloudy atmosphere in general.

The Fourier analysis of the measured time series of cloud top height shows that the calculated slopes of the power spectrum are with values of 1.8–2.0 larger than the widely used assumption of 5/3. So the spatial small-scale variability seems to be not so important to describe the variance of a cloud top.

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The examined cases show an increase of the differences between homogeneous and heterogeneous albedo values at larger cloud top variability. With it convective clouds with often larger geometrical cloud top variability are more affected by the influence of these heterogeneity effects than stratiform clouds for example. The solar zenith angle has a larger influence on radiative transfer than cloud top variability, but for large solar zenith angles the differences of albedo values are negligible.

The differences of reflectances are also larger for higher solar zenith angles, so the negative maximum (homogeneous > heterogeneous) appears in backward direction and the positive maximum in forward direction. The largest difference of the mean reflectances between homogeneous and heterogeneous cloud top is approximately 0.3, so it can be important for measurements in these directions. Many atmospheric parameters are deviated from satellite measurements. The radiances are gauged at the VIS and IR spectra. The influence of cloud top variability on radiative transfer is not the most important one of course, but might have influence on the accuracy of deviated parameters.

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Table 1. Cloudfield parameters.

	cloud 1	cloud 2	cloud 3
CB [m]	2200	1500	2800
mean CT [m]	3775.4	3623.3	3382.0
std. deviation CT [m]	135.4	178.8	132.1
slope CT	1.987	1.832	2.009
temperature CT [°K]	263	263	265
temperature CB [°K]	270	278	269
pressure CB [hPa]	783.5	852.5	722.5

Table 2. Calculated albedo values [%] for the different cloud fields and solar zenith angle (θ_{sun}), the difference is defined as heterogeneous minus homogeneous albedo value.

θ_{sun}	0°	30°	60°
cloud 1			
homogeneous	10.568	10.640	10.826
heterogeneous	10.339	10.590	10.841
difference	-0.229	-0.050	0.015
cloud 2			
homogeneous	10.877	10.912	11.020
heterogeneous	10.785	10.896	11.032
difference	-0.092	-0.016	0.012
cloud 3			
homogeneous	8.587	8.817	9.472
heterogeneous	7.456	8.685	9.478
difference	-1.131	-0.132	0.006

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Table 3. Transmission and absorption [%] for the different cloud fields (described in Table 1) and solar zenith angles.

variability		θ_{sun}		
		0°	30°	60°
transmission				
cloud 1	homogeneous	6.6920	6.1150	4.5183
	heterogeneous	8.8190	6.5252	4.2626
cloud 2	homogeneous	3.8751	3.5504	2.6138
	heterogeneous	4.5957	3.6744	2.5222
cloud 3	homogeneous	24.9300	22.8276	16.8455
	heterogeneous	35.1987	24.0418	16.8273
absorption				
cloud 1	homogeneous	0.1385	0.1276	0.0963
	heterogeneous	0.1705	0.1329	0.0937
cloud 2	homogeneous	0.2619	0.2409	0.1795
	heterogeneous	0.3039	0.2471	0.1750
cloud 3	homogeneous	0.0203	0.0192	0.0155
	heterogeneous	0.0219	0.0199	0.0162

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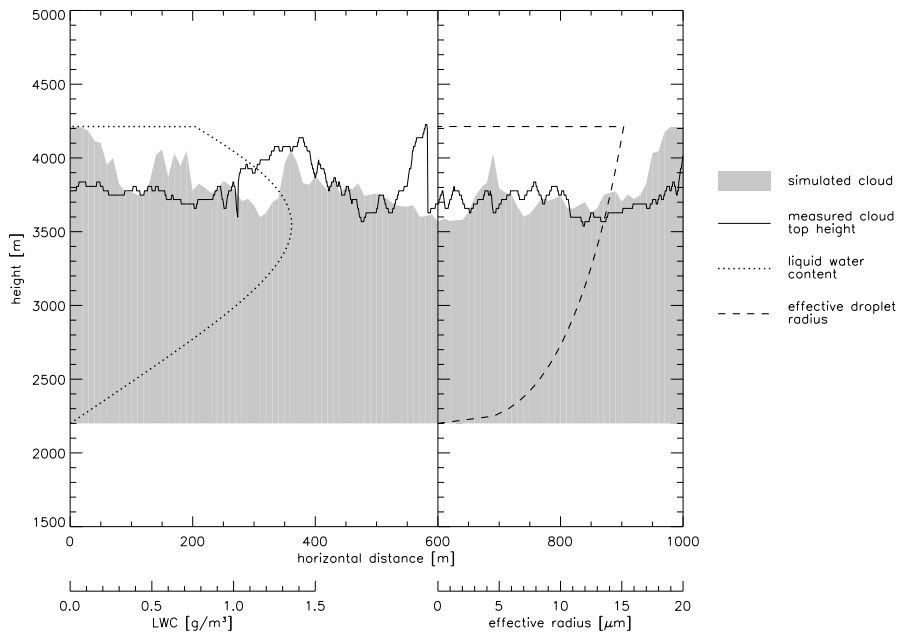


Fig. 1. Measured and simulated cloud top height, subadiabatic LWC profile and adiabatic profile of the effective radius.

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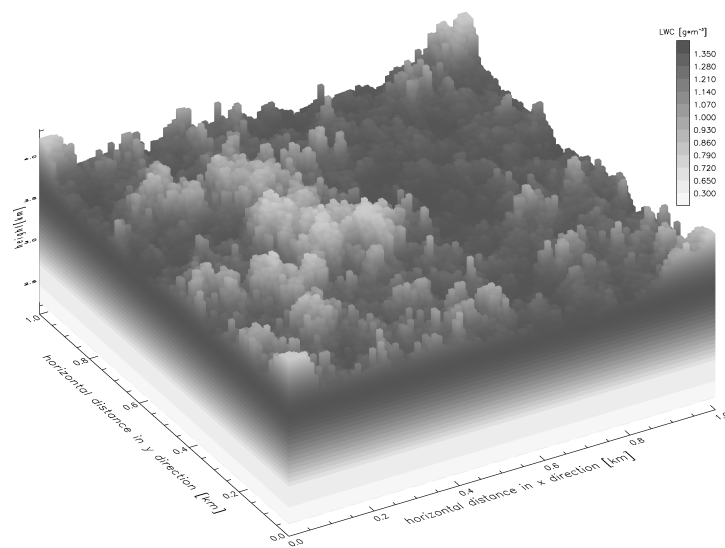


Fig. 2. LWC of the 3-D simulated cloud field.

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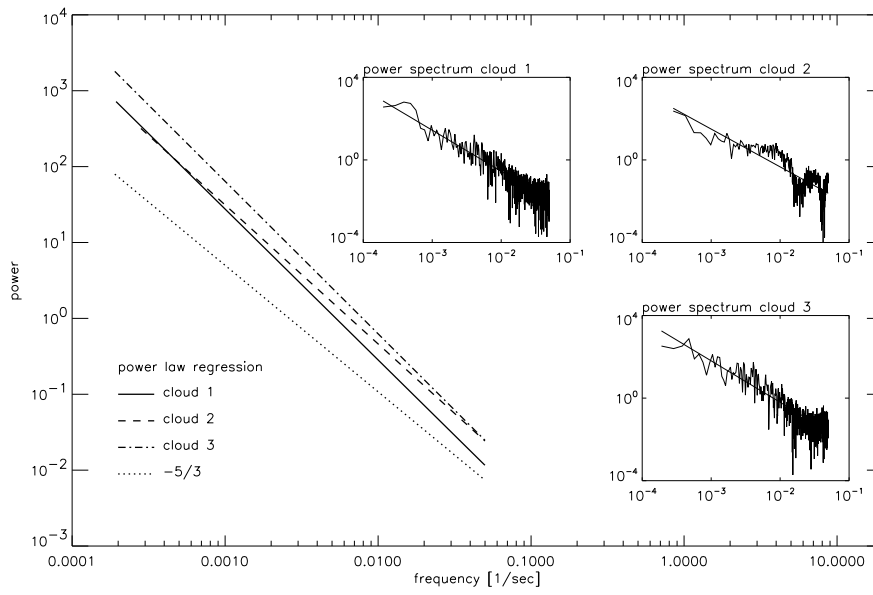


Fig. 3. Comparison of calculated power spectra.

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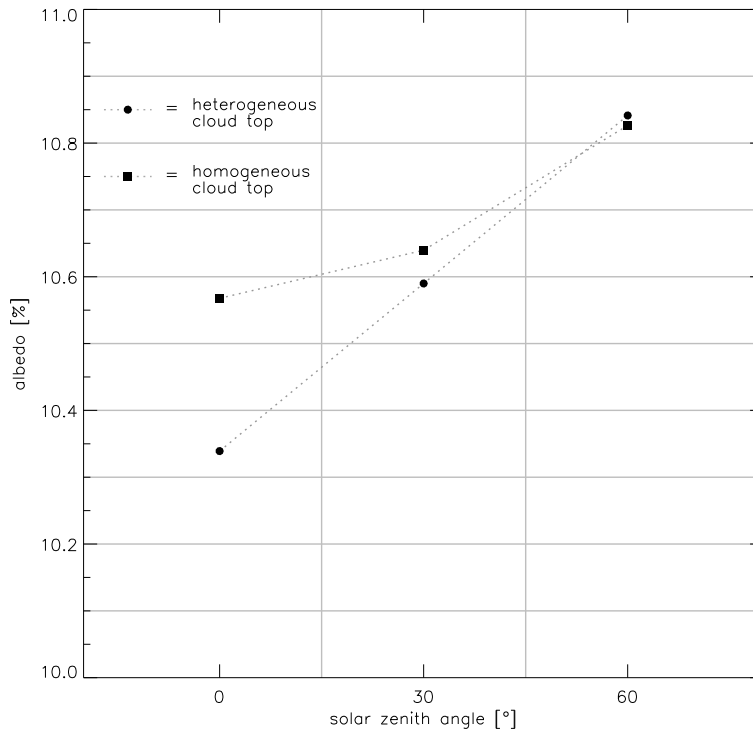


Fig. 4. Mean albedo values for cloud 1.

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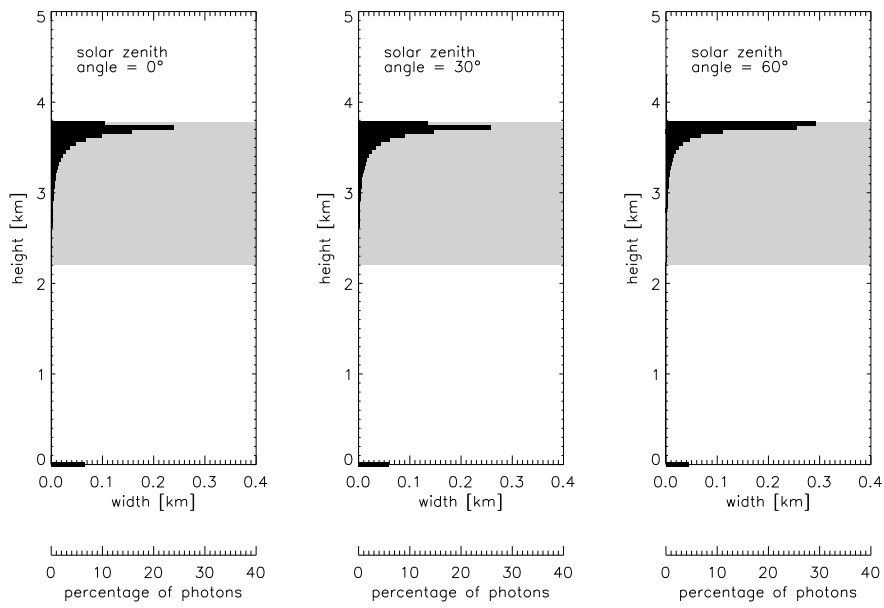


Fig. 5. Penetration depth of cloud 1 (homogeneous top) for different solar zenith angles.

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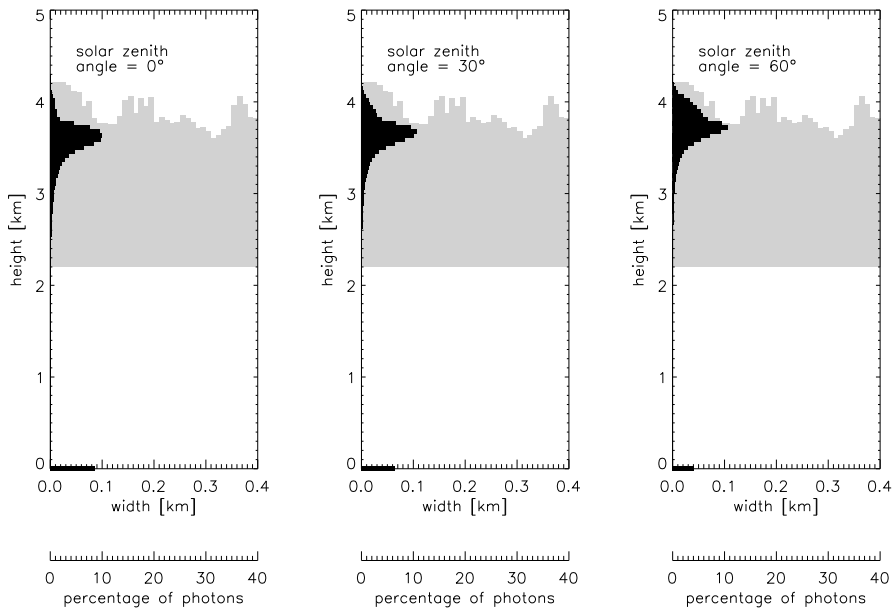


Fig. 6. Penetration depth of cloud 1 (heterogeneous top) for different solar zenith angles.

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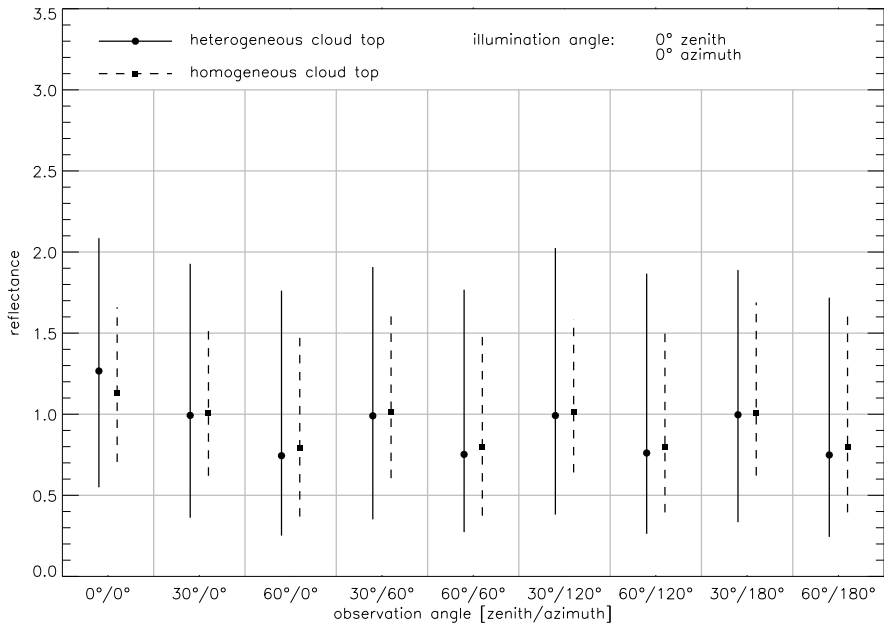


Fig. 7. Reflectances of cloud 1 ($\theta_{\text{sun}} = 0^\circ$) for different observation angles.

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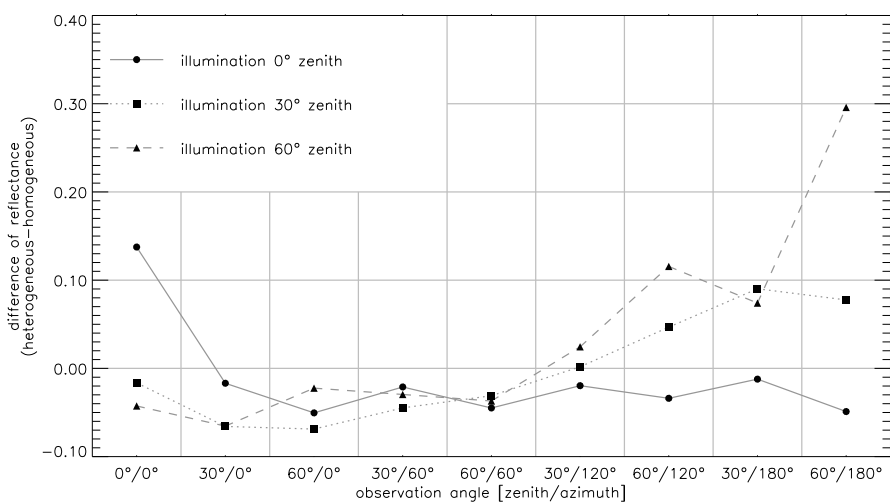


Fig. 8. Differences between heterogeneous and homogeneous reflectances for several illumination and observation angles.

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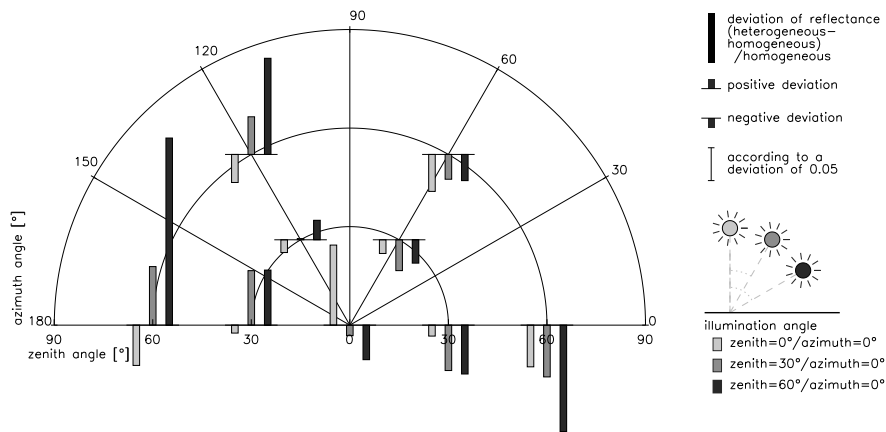


Fig. 9. Deviation of reflectance ((heterogeneous – homogeneous)/homogeneous) for several illumination and observation angles.